

# Lesson 5 - Relativistic Wave Equations

## Unit 5.1 Klein-Gordon and Dirac equations

Luca Salasnich

Dipartimento di Fisica e Astronomia "Galileo Galilei", Università di Padova

Structure of Matter - MSc in Physics

# Klein-Gordon equation (I)

The classical energy of a nonrelativistic free particle is given by

$$E = \frac{\mathbf{p}^2}{2m}, \quad (1)$$

where  $\mathbf{p}$  is the linear momentum and  $m$  the mass of the particle. The Schrödinger equation of the corresponding quantum particle with wavefunction  $\psi(\mathbf{r}, t)$  is easily obtained by imposing the quantization prescription

$$E \rightarrow i\hbar \frac{\partial}{\partial t}, \quad \mathbf{p} \rightarrow -i\hbar \nabla. \quad (2)$$

In this way one gets the time-dependent Schrödinger equation of the free particle, namely

$$i\hbar \frac{\partial}{\partial t} \psi(\mathbf{r}, t) = -\frac{\hbar^2}{2m} \nabla^2 \psi(\mathbf{r}, t), \quad (3)$$

obtained for the first time in 1926 by Erwin Schrödinger.

## Klein-Gordon equation (II)

The classical energy of a relativistic free particle is instead given by

$$E = \sqrt{\mathbf{p}^2 c^2 + m^2 c^4} , \quad (4)$$

where  $c$  is the speed of light in the vacuum. By applying directly the quantization prescription (2) one finds

$$i\hbar \frac{\partial}{\partial t} \psi(\mathbf{r}, t) = \sqrt{-\hbar^2 c^2 \nabla^2 + m^2 c^4} \psi(\mathbf{r}, t) . \quad (5)$$

This equation is quite suggestive but the square-root operator on the right side is a very difficult mathematical object.

For this reason in 1927 Oskar Klein and Walter Gordon suggested to start with

$$E^2 = \mathbf{p}^2 c^2 + m^2 c^4 \quad (6)$$

and then to apply the quantization prescription (2). In this way one obtains

$$\left( \frac{1}{c^2} \frac{\partial^2}{\partial t^2} - \nabla^2 + \frac{m^2 c^2}{\hbar^2} \right) \psi(\mathbf{r}, t) = 0 , \quad (7)$$

i.e. a generalization of Maxwell's wave equation for massive particles.

# Klein-Gordon equation (III)

The Klein-Gordon equation has two problems:

- i) it admits solutions with negative energy;
- ii) the space integral over the entire space of the non negative probability density  $\rho(\mathbf{r}, t) = |\psi(\mathbf{r}, t)|^2 \geq 0$  is generally not time-independent, namely

$$\frac{d}{dt} \int_{\mathbb{R}^3} \rho(\mathbf{r}, t) d^3\mathbf{r} \neq 0. \quad (8)$$

Nowadays we know that to solve completely these two problems it is necessary to promote  $\psi(\mathbf{r}, t)$  to a quantum field operator. Within this second-quantization (quantum field theory) approach the Klein-Gordon equation is now used to describe relativistic particles with spin zero, like the pions or the Higgs boson.

# Dirac equation (I)

In 1928 Paul Dirac proposed a different approach to the quantization of the relativistic particle. To solve the problem of Eq. (8) he considered a wave equation with only first derivatives with respect to time and space and introduced the classical energy

$$E = c \hat{\alpha} \cdot \mathbf{p} + \hat{\beta} mc^2, \quad (9)$$

such that squaring it one recovers the Klein-Gordon equation. This condition is fulfilled only if  $\hat{\alpha} = (\hat{\alpha}_1, \hat{\alpha}_2, \hat{\alpha}_3)$  and  $\hat{\beta}$  satisfy the following algebra of matrices

$$\hat{\alpha}_1^2 = \hat{\alpha}_2^2 = \hat{\alpha}_3^2 = \hat{\beta}^2 = \hat{1}, \quad (10)$$

$$\hat{\alpha}_i \hat{\alpha}_j + \hat{\alpha}_j \hat{\alpha}_i = \hat{0}, \quad i \neq j \quad (11)$$

$$\hat{\alpha}_i \hat{\beta} + \hat{\beta} \hat{\alpha}_i = \hat{0}, \quad \forall i \quad (12)$$

where  $\hat{1}$  is the identity matrix and  $\hat{0}$  is the zero matrix.

## Dirac equation (II)

The smallest dimension in which the so-called Dirac matrices  $\hat{\alpha}_i$  and  $\hat{\beta}$  can be realized is four. In particular, one can write

$$\hat{\alpha}_i = \begin{pmatrix} \hat{0}_2 & \hat{\sigma}_i \\ \hat{\sigma}_i & \hat{0}_2 \end{pmatrix}, \quad \hat{\beta} = \begin{pmatrix} \hat{I}_2 & \hat{0}_2 \\ \hat{0}_2 & -\hat{I}_2 \end{pmatrix}, \quad (13)$$

where  $\hat{I}_2$  is the  $2 \times 2$  identity matrix,  $\hat{0}_2$  is the  $2 \times 2$  zero matrix, and

$$\hat{\sigma}_1 = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} \quad \hat{\sigma}_2 = \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix} \quad \hat{\sigma}_3 = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} \quad (14)$$

are the Pauli matrices. Eq. (9) with the quantization prescription (2) gives

$$i\hbar \frac{\partial}{\partial t} \Psi(\mathbf{r}, t) = \left( -i\hbar c \hat{\alpha} \cdot \nabla + \hat{\beta} mc^2 \right) \Psi(\mathbf{r}, t), \quad (15)$$

which is the Dirac equation for a free particle.

# Dirac equation (III)

The wavefunction  $\Psi(\mathbf{r}, t)$  of the Dirac equation has four components in the abstract space of Dirac matrices, i.e. this spinor field can be written

$$\Psi(\mathbf{r}, t) = \begin{pmatrix} \psi_1(\mathbf{r}, t) \\ \psi_2(\mathbf{r}, t) \\ \psi_3(\mathbf{r}, t) \\ \psi_4(\mathbf{r}, t) \end{pmatrix}. \quad (16)$$

In explicit matrix form the Dirac equation is thus given by

$$i\hbar \frac{\partial}{\partial t} \begin{pmatrix} \psi_1(\mathbf{r}, t) \\ \psi_2(\mathbf{r}, t) \\ \psi_3(\mathbf{r}, t) \\ \psi_4(\mathbf{r}, t) \end{pmatrix} = \hat{H} \begin{pmatrix} \psi_1(\mathbf{r}, t) \\ \psi_2(\mathbf{r}, t) \\ \psi_3(\mathbf{r}, t) \\ \psi_4(\mathbf{r}, t) \end{pmatrix} \quad (17)$$

where the matrix operator  $\hat{H}$  is given by

$$\begin{pmatrix} mc^2 & 0 & -i\hbar c \frac{\partial}{\partial z} & -i\hbar c \left( \frac{\partial}{\partial x} - i \frac{\partial}{\partial y} \right) \\ 0 & mc^2 & -i\hbar c \left( \frac{\partial}{\partial x} + i \frac{\partial}{\partial y} \right) & i\hbar c \frac{\partial}{\partial z} \\ -i\hbar c \frac{\partial}{\partial z} & -i\hbar c \left( \frac{\partial}{\partial x} - i \frac{\partial}{\partial y} \right) & -mc^2 & 0 \\ -i\hbar c \left( \frac{\partial}{\partial x} + i \frac{\partial}{\partial y} \right) & i\hbar c \frac{\partial}{\partial z} & 0 & -mc^2 \end{pmatrix}.$$

## Dirac equation (IV)

It is easy to show that the Dirac equation satisfies the differential law of current conservation, given by

$$\frac{\partial}{\partial t} \rho(\mathbf{r}, t) + \nabla \cdot \mathbf{j}(\mathbf{r}, t) = 0, \quad (19)$$

where

$$\rho(\mathbf{r}, t) = \Psi^\dagger(\mathbf{r}, t) \Psi(\mathbf{r}, t) = \sum_{i=1}^4 |\psi_i(\mathbf{r}, t)|^2 \quad (20)$$

is the probability density, and  $\mathbf{j}(\mathbf{r}, t)$  is the probability current with three components

$$j_k(\mathbf{r}, t) = c \Psi^\dagger(\mathbf{r}, t) \hat{\alpha}_k \Psi(\mathbf{r}, t). \quad (21)$$

Finally, we observe that from the continuity equation (19) one finds

$$\frac{d}{dt} \int_{\mathbb{R}^3} \rho(\mathbf{r}, t) d^3\mathbf{r} = 0, \quad (22)$$

by using the divergence theorem and imposing a vanishing current density on the border at infinity. Thus, contrary to the Klein-Gordon equation, the Dirac equation does not have the probability density problem.